

LLST.1 | 1/2/16.1 | Liverpool. Lect. 1 | Why string?

physics: Mathematical formulation of experimental observations

unification: encompassing mathematical model.

strings: unification of:

Gravity

Celestial & cosmological

Riemannian Geometry

General relativity

Gauge & Matter.

subatomic particles & fields

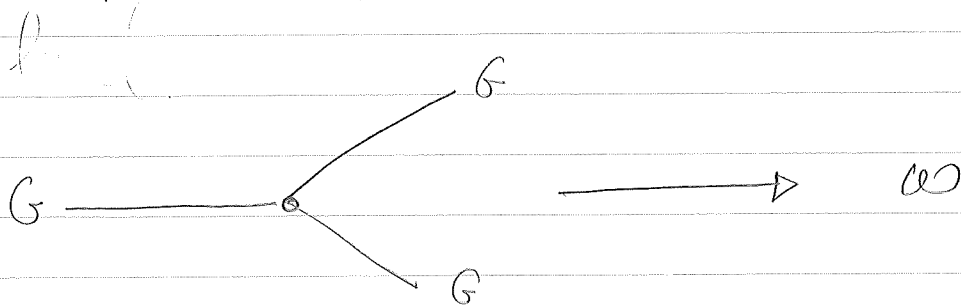
Quantum field theory.

Standard Model.

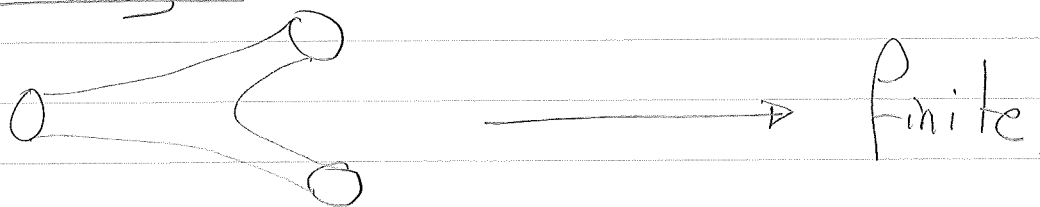
yet.

General relativity & QFT's are

incompatible at a fundamental level.



Strings

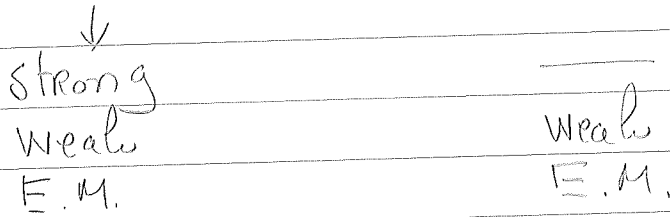


LLST.2 | 2/2/16.1 | Liverpool

Inventory

Forces	E.M. x Weak	Strong	spin +1 particles
Symmetries	U(1) x SU(2)	SU(3)	local phase invariance
	Gravity		spin 2 particle

Matter: Quarks & Leptons.



quarks: $\begin{pmatrix} \text{up} \\ \text{down} \end{pmatrix}$ $\begin{pmatrix} c \\ s \end{pmatrix}$ $\begin{pmatrix} t \\ b \end{pmatrix}$

leptons: $\begin{pmatrix} \nu_e \\ e \end{pmatrix}$ $\begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}$ $\begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix}$

mass \rightarrow spontaneous symmetry breaking \rightarrow Higgs

$SU(2) \times U(1) \rightarrow U(1)_{\text{e.m.}}$

Unification: GUTs

$SU(3) \times SU(2) \times U(1) \rightarrow SU(5) \rightarrow SO(10)$

SO \rightarrow simple orthogonal 10×10 matrices: generators $\sigma^T = -C$

$\begin{pmatrix} \square \\ \diagdown \end{pmatrix} \frac{n^2-n}{2} + n = \frac{n(n+1)}{2} \rightarrow$ symmetric.

$\begin{pmatrix} \square \\ \diagup \end{pmatrix} \frac{n^2-n}{2} = \frac{n(n-1)}{2} \rightarrow$ asymmetric (i) hermitian $\sigma^\dagger = 0$

LLST, 3 | 27/1/14, 1 | LIVERPOOL (revised).

Time evolution: $|a(0)\rangle \rightarrow |a(t)\rangle$

$$|a(t)\rangle = \bar{U} |a(0)\rangle$$

Physical theories \rightarrow Unitary $U^\dagger U = 1$

$$\Rightarrow \langle a(t) | a(t) \rangle = \langle a(0) | U^\dagger U | a(0) \rangle = \langle a(0) | a(0) \rangle$$

Representation of a unitary operator:

$$U = e^{iHt} \leftarrow \text{hermitian}$$

$H \rightarrow$ generators of a Lie algebra

eg. $SU(n)$, $SO(2n)$.

$$n=5 \quad \begin{matrix} & & 2 & & 1 & & 1 & & 1 \\ SO(10) & \rightarrow & SU(3) & \times & SU(2) & \times & U(1)_x & \times & U(1)_y \end{matrix}$$

$$\begin{aligned} 45 = & (8, 1, 0, 0) + (1, 3, 0, 0) + (3, 2, x, y) + (\bar{3}, 2, -x, -y) \\ & + (6, 1, x', y') + (\bar{6}, 1, -x', -y') \\ & + (1, 3, x'', y'') + (1, 3, -x'', -y'') \\ & + (1, 1, x''', y''') + (1, 1, -x''', -y'''). \end{aligned}$$

LLST.3 | 27/1/14.2 | LIVERPOOL / REVISÉ

FOR OUR PURPOSE: chiral 16 representation of $SO(10)$ put even - in 5 slots

3			2		
+	+	+	+	+	e_L^c
+	+	+	-	-	N_L^c
+	+	-	+	-	
+	-	+	+	-	
-	+	+	+	-	$Q = \begin{pmatrix} u \\ d \end{pmatrix}$
+	+	-	-	+	
+	-	+	-	+	
-	+	+	-	+	
-	-	+	+	+	
-	+	-	+	+	D_L^c
+	-	-	+	+	
-	-	-	-	+	L
-	-	-	+	-	
-	-	+	-	-	
-	+	-	-	-	U_L^c
+	-	-	-	-	

LLST.4 | 26/6/09.4 | Abmgdon.

OR - using $\binom{n}{k} = \frac{n!}{k!(n-k)!}$

$$16 = \binom{3}{0} \binom{2}{0} + \binom{3}{0} \binom{2}{2} + \binom{3}{1} \binom{2}{1} + \binom{3}{2} \binom{2}{0} + \binom{3}{2} \binom{2}{2} + \binom{3}{3} \binom{2}{1}$$

$(1, 1) \quad (1, 1) \quad (3, 2) \quad (\bar{3}, 1) \quad (\bar{3}, 1) \quad (1, 2)$

$E_L^c \quad N_L^c \quad Q_L \quad D_L^c \quad U_L^c \quad L_L$

These are the sixteen states in a one S, M_0 generation.
+ R. H. N.

There is something right about GUTs (gauged)

Problems | Hierarchy? $M_W \sim 100 \text{ GeV}$ } ?
 $M_{\text{GUT}} \sim 10^{16} \text{ GeV}$ }

Flavour? Not one but three?

GUTs: gauge theories \rightarrow renormalizable ^{local} quantum field theories

Renormalizable: infinities absorbed in a finite number
of measurable parameters. \rightarrow

Gravity \rightarrow nonrenormalizable.

\rightarrow Quantum of Gravity?

\rightarrow String \rightarrow finite \rightarrow Gravity & Gauge unification.

LLST.5 / 3/7/09.1 / Abington / Ridley /

special relativity and extra dimensions

Einstein's special relativity: 1) c = speed of light = constant in all inertial frames

$$c = 3 \times 10^8 \frac{m}{s}$$

2) Laws of physics are the same in all inertial frames.

events are labeled by their time & position in

Inertial frames. \rightarrow 4-vector: $X^\mu = (ct, \vec{X})$
 $X^0 = ct$ $\vec{X} = (x^1, x^2, x^3) = (x, y, z)$ $\mu = 0, 1, 2, 3$

The length of four vector is given by:

$$X \cdot X = -c^2 t^2 + x^2 + y^2 + z^2$$

$$= \sum_{\mu, \nu=0}^3 \eta_{\mu\nu} X^\mu X^\nu = X_\mu X^\mu$$

Einstein summation convention.

where

$$\eta_{\mu\nu} = \begin{pmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{pmatrix} \begin{matrix} \leftarrow \text{time} \\ \left. \begin{matrix} \\ \\ \end{matrix} \right\} \text{3 space} \end{matrix}$$

is the Minkowski metric.

In general the scalar product of two four vectors

X^μ, Y^ν is given by $X \cdot Y = \sum_{\mu, \nu=0}^3 \eta_{\mu\nu} X^\mu Y^\nu = X_\nu Y^\nu$
 $[X_\nu = \eta_{\mu\nu} X^\mu]$

Lorentz Transformations:

From one inertial frame X^μ
To another " " X'^μ .

The Lorentz transformations preserve the scalar product.

$$-c^2 t^2 + \vec{x}^2 = X \cdot X = \eta_{\mu\nu} X^\mu X^\nu = X' \cdot X' = \eta_{\mu\nu} X'^\mu X'^\nu = -c^2 t'^2 + \vec{x}'^2$$

where X and X' are related by a Lorentz transformation.

$$X^\mu \rightarrow X'^\mu = \Lambda^\mu{}_\nu X^\nu$$

Think for simplicity with the more familiar case of ordinary rotations in 3 dimensions.

$$\vec{x} \rightarrow \vec{y} = R \vec{x}$$

Rotations preserve the length of a three vector similarly; Lorentz transformations preserve the length of a four vector in 4 dimensional Minkowski space-time.

*** Properties of Lorentz transformations

$\eta_{\mu\nu} X^\mu X^\nu \rightarrow$ invariant under L.T.

No change in size or shape

$$\begin{aligned} \eta_{\mu\nu} X^\mu X^\nu &\rightarrow \eta_{\mu\nu} X'^\mu X'^\nu = \eta_{\mu\nu} \Lambda^\mu{}_\alpha X^\alpha \Lambda^\nu{}_\beta X^\beta \\ &= \eta_{\alpha\beta} X^\alpha X^\beta \end{aligned}$$

LLST.7 | 3/7/09.3 | Abingdon / Friday |

$$\Rightarrow (*) \quad \eta_{\mu\nu} \Lambda^\mu_\alpha \Lambda^\nu_\beta = \eta_{\alpha\beta}$$

In matrix notation $\Lambda^T \eta \Lambda = \eta$.

→ Defines the Lorentz transformations.

$$\Rightarrow |\det \Lambda|^2 = 1 \Rightarrow \det \Lambda = \pm 1.$$

The physical case $\det \Lambda = +1$.

$\det \Lambda = 1 \rightarrow$ Proper Lorentz transformations.

$\det \Lambda = -1 \rightarrow$ Improper Lorentz " "

Look at 00 component of $\Lambda^\mu_\alpha \Lambda^\nu_\beta \eta_{\mu\nu} = \eta_{\alpha\beta}$.

$$\eta_{\mu\nu} \Lambda^\mu_0 \Lambda^\nu_0 = \eta_{00} = -1$$
$$-(\Lambda^0_0)^2 + \sum_{i=1}^3 (\Lambda^i_0)^2 = -1$$

$$(\Lambda^0_0)^2 = 1 + \sum_i (\Lambda^i_0)^2 \geq +1$$

$$(\Lambda^0_0)^2 \geq +1 \begin{cases} \Lambda^0_0 \geq +1 & \text{orthochronous} \\ \Lambda^0_0 \leq -1 & \text{non-orthochronous} \end{cases}$$

Examples of Lorentz transformations

1) $\Lambda^\mu_\nu = \begin{pmatrix} -1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix} \rightarrow$ Proper - non-orthochronous
reflection.

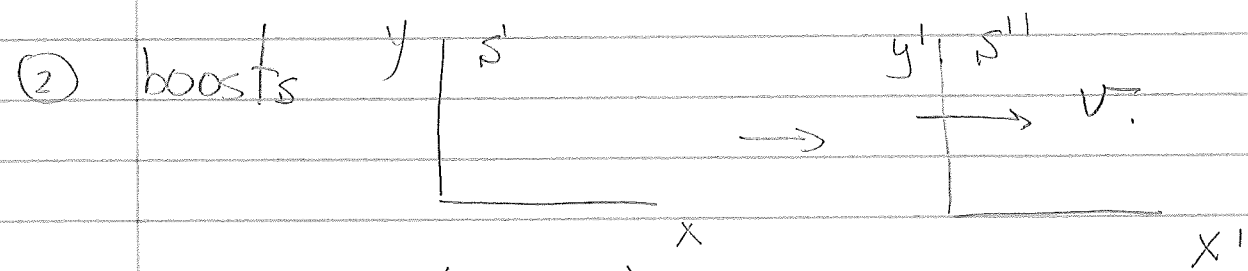
2) $\Lambda_{\nu}^{\mu} = \begin{pmatrix} -1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$ $t \rightarrow -t$ $\vec{x} \rightarrow \vec{x}$
 \rightarrow improper \rightarrow non-orthochronous

3) Physical Lorentz transformations \rightarrow
 \rightarrow continuously connected to the identity \rightarrow

- \rightarrow (a) $\det \Lambda = 1 \leftarrow$ Proper.
- (b) $\Lambda^0_0 \geq 1 \leftarrow$ orthochronous.

Examples | $\Lambda_{\nu}^{\mu} = \begin{pmatrix} 1 & & & \\ & 1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$ $\Lambda_{\nu}^{\mu} = \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & 1 & \\ & & & 1 \end{pmatrix}$

(1) Rotations $\Lambda = \begin{pmatrix} 1 & 0 \\ 0 & R \end{pmatrix}$ $\det \Lambda = \det R$
 $\det R = \pm 1$
 $\det R = 1 \rightarrow$ proper.



$ct' = \gamma(ct - \beta x)$ $\gamma = \frac{1}{\sqrt{1 - \beta^2}}$ $\beta = \frac{v}{c}$
 $x' = \gamma(x - \beta ct)$

$y' = y$
 $z' = z$

OR

$X'^0 = \gamma(X^0 - \beta X^1)$
 $X'^1 = \gamma(-\beta X^0 + X^1)$
 $X'^2 = X^2$
 $X'^3 = X^3$

LLST, 9 | 3/7/09.5 | Abingdon | Friday

$$\text{OR} \quad \Lambda^\mu = \begin{pmatrix} \gamma & -\gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix}$$

$$\det \Lambda = \gamma^2 - \gamma^2 \beta^2 = \frac{1}{|1 - \beta^2|} (1 - \beta^2) = \underline{1}$$

$$\Lambda^0_0 = \gamma = \frac{1}{\sqrt{1 - \beta^2}} \geq 1 \quad \text{as } \beta = \frac{v}{c} < 1.$$

The Poincare group

consider the infinitesimal relativistic line element

$$(*) \quad -ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2.$$

so far we discussed rotation & boosts
which are the Lorentz transformations.

The most general transformations also include
translations. $X^\mu \rightarrow X'^\mu = X^\mu + a^\mu$

where a^μ is a constant 4-vector

$$\text{obviously} \quad dt' = dt \quad dx'^i = dx^i.$$

we have:

4 - translations	dt, dx, dy, dz
3 - boosts	$dt dx, dt dy, dt dz$
3 - rotations	$dx dy, dx dz, dy dz$

These are the Poincare transformations.

→ They form the Poincare group.

LST, 10 | 3/7/09.6 | Abingdon | Friday

Representation of the Poincare group are labeled by:

- 1) spin \rightarrow label of the Lorentz group.
- 2) mass

spin	Examples	spin	components	
		0	1	scalar
		$\frac{1}{2}$	2	Weyl spinor
		$\frac{1}{2}$	4	Dirac spinor = W-spinor + W-spinor
		+1	4	vector

Relativistic energy and momentum

Non relativistic energy-momentum relation: $E = \frac{\vec{p}^2}{2m} + \frac{mv^2}{2}$

relativistic " " " $\frac{E^2}{c^2} = \vec{p} \cdot \vec{p} + m^2 c^2$

relativistic $E = \gamma m c^2$ $\vec{p} = \gamma m \vec{v}$

Energy-momentum 4-vector $P^\mu = (p^0, p^1, p^2, p^3) = (\frac{E}{c}, p^1, p^2, p^3)$
 $= (\frac{E}{c}, \vec{p}) = m\gamma(c, \vec{v})$

$$\Rightarrow P_\mu P^\mu = \eta_{\mu\nu} P^\mu P^\nu = -\frac{E^2}{c^2} + \vec{p} \cdot \vec{p} = -m^2 c^2 = P^2$$

$P_\mu P^\mu = -m^2 c^2 \rightarrow$ Lorentz scalar (invariant)

time intervals are not Lorentz invariant.

Proper time - time measured in the moving-frame relative to itself.

LST.11 | 5/7/09.1 | Abingdon | Sunday

consider a particle moving along the x -axis.

consider: $-ds^2 = -c^2 dt^2 + dx^2 = -c^2 dt^2 (1 - \beta^2)$. ($dx = \beta c dt$)

evaluate the interval in a frame moving with the particle

$$dx' = 0 \quad dt' = dt_p \rightarrow \text{proper time.}$$

$$ds^2 = c^2 dt_p^2 \Rightarrow ds = c dt_p \Rightarrow \frac{ds}{c} = dt_p \quad \uparrow \text{proper time}$$

also $ds = c dt (\sqrt{1 - \beta^2}) \Rightarrow \frac{dt}{ds} = \frac{\gamma}{c}$

$ds \rightarrow$ Lorentz scalar \rightarrow construct new Lorentz vectors

velocity four vector: $u^\mu = c \frac{dx^\mu}{ds} = c \left(\frac{d(ct)}{ds}, \frac{dx}{ds}, \frac{dy}{ds}, \frac{dz}{ds} \right)$

simplify e.g. $\frac{dx}{ds} = \left(\frac{dx}{dt} \right) \left(\frac{dt}{ds} \right) = v_x \frac{\gamma}{c}$

$$\Rightarrow u^\mu = \gamma (c, v_x, v_y, v_z)$$

$$\Rightarrow p^\mu = m u^\mu$$

any Lorentz 4-vector transforms under

Lorentz transformations as X^μ ,

$$P'^\mu = \Lambda^\mu_\nu P^\nu$$

For x -boost

$$\begin{pmatrix} E'/c \\ P'_x \\ P'_y \\ P'_z \end{pmatrix} = \begin{pmatrix} \gamma - \gamma\beta & 0 & 0 \\ -\gamma\beta & \gamma & 0 & 0 \\ 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} E/c \\ P_x \\ P_y \\ P_z \end{pmatrix} = \begin{pmatrix} \gamma (E/c - \beta P_x) \\ \gamma (-\beta E/c + P_x) \\ P_y \\ P_z \end{pmatrix}$$

LST.12 | 5/7/09, 2 | Abingdon | Sunday |

Lorentz invariance in extra dimensions

$$-ds^2 = -c^2 dt^2 + dx^2 + dy^2 + dz^2 + dw^2 + \dots$$

$$\eta_{\mu\nu} = \begin{pmatrix} -1 & & & & & \\ & +1 & & & & \\ & & +1 & & & \\ & & & +1 & & \\ & & & & +1 & \\ & & & & & +1 \end{pmatrix}$$

etc...

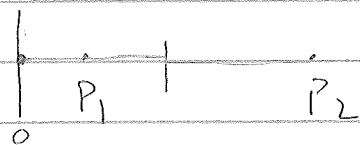
more components but otherwise the same.

compact extra dimensions.

only 4- are observed.

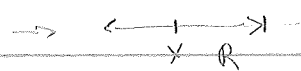
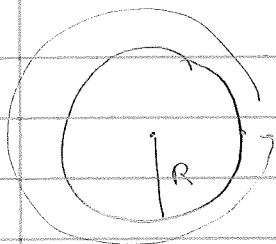
⇒ Extra dimensions are compact - curled.

x_4



$$x(P_2) = x(P_1) + 2\pi R n,$$

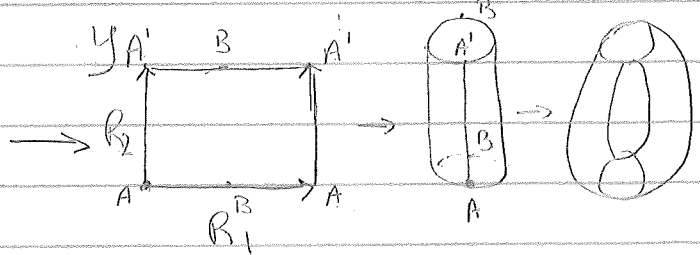
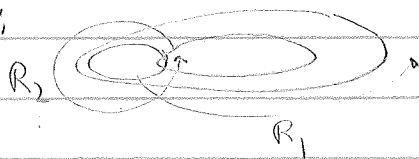
n - integer R - constant.



$$x \sim x + 2\pi R n, \leftarrow \text{circle.}$$

$0 \leq x < 2\pi R$ - fundamental domain.

Torus:



$$x \sim x + 2\pi R_1 n$$

$$y \sim y + 2\pi R_2 n$$

LST. 13 | 5/8/09.1 | Abingdon / Wednesday |

consider the real line $-\infty \leq x \leq \infty$

impose the condition $x \sim -x$.

points are identified under the \mathbb{Z}_2 reflection.

The fundamental domain is now $x \geq 0$.

$x=0$ is a fixed point.

This is the orbifold $\mathbb{R}_1 / \mathbb{Z}_2$

The orbifold is singular at the fixed point.

Quantum mechanics and the square well

Planck constant expresses the relation between the energy of a photon and its angular frequency.

$$E = \hbar \omega.$$

it is the fundamental constant of quantum mechanics

it appears in the basic commutation relations of quantum mechanics.

$$[\hat{x}, \hat{p}] = i\hbar.$$

and the limit $\lim_{\hbar \rightarrow 0}$ gives the classical limit.

From $E = \hbar \omega$ its units are

$$[E] = \frac{[M] \cdot [L]^2}{[T]^2} = [\hbar] \cdot \frac{1}{[T]} \Rightarrow [\hbar] = \frac{[M] [L]^2}{[T]} = [E][T] = [S]$$

$m v^2$

[action] in Lagrangian mechanics

LST. 14 | 5/8/09.2 | Abingdon | Wednesday |

In Extra dimensions.

$$[X^i, P_j] = i\hbar \delta^i_j.$$

where

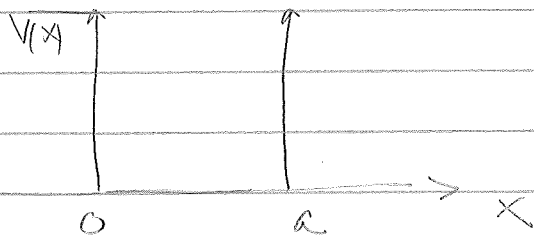
$$\delta^i_j = \begin{cases} 1 & i=j \\ 0 & i \neq j \end{cases}$$

consider the time-independent Schrödinger eq. in 1-dimension.

$$-\frac{\hbar^2}{2m} \nabla^2 \psi(x) + V(x) \psi(x) = E \psi(x)$$

in the case of a square-well potential of infinite height.

$$V(x) = \begin{cases} 0 & x \in (0, a) \\ \infty & x \notin (0, a) \end{cases}$$



The boundary conditions imply that

$$\psi(x) = 0 \quad \text{for } x \notin (0, a)$$

continuity of the wave function, on the

boundary implies that $\psi(0) = \psi(a) = 0$.

\Rightarrow quantum mechanics of a particle confined to the segment $0 \leq x \leq a$.

LST, 15 | 5/8/09.3 | Abingdon | Wednesday |

$$\text{For } 0 \leq x \leq a. \quad -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} = E\psi \quad (*)$$

$$\Rightarrow \psi(x) = A \sin \lambda x + B \cos \lambda x.$$

$$\psi(0) = 0 \Rightarrow B = 0.$$

$$\psi(a) = 0 \Rightarrow A \sin \lambda a = 0 \Rightarrow \lambda = \frac{k\pi}{a}$$

$$k = 1, 2, 3, \dots, \infty$$

$$\text{sub into } (*) \Rightarrow +\frac{\hbar^2}{2m} \frac{k^2 \pi^2}{a^2} A \sin \frac{k\pi}{a} x = E A \sin \frac{k\pi}{a} x$$

$$\Rightarrow E = \frac{\hbar^2 k^2 \pi^2}{2ma^2}$$

$$\text{To find } A \text{ impose } \int_{-a}^a |\psi|^2 dx = 1 = \int_{-a}^a \psi^2 dx$$

$$\begin{aligned} \cos^2 - \sin^2 &= \cos 2x \\ \cos^2 + \sin^2 &= 1 \end{aligned} = \int_0^a A^2 \frac{\sin^2 \frac{k\pi}{a} x}{a} dx = A^2 \int_0^a \frac{1}{2} \left(1 - \cos 2 \frac{k\pi}{a} x \right) dx =$$

$$= \frac{A^2}{2} \left(x \Big|_0^a - \frac{a}{2k\pi} \sin 2 \frac{k\pi}{a} x \Big|_0^a \right) = \frac{A^2 a}{2} = 1$$

$$\Rightarrow A = \sqrt{\frac{2}{a}} \Rightarrow \psi(x) = \sqrt{\frac{2}{a}} \sin \frac{k\pi x}{a}$$

square well with extra dimension.

consider the case with an extra compact dimension.

$$(x, y) \sim (x, y + 2\pi R)$$

with

$$V(x, y) = \begin{cases} \infty & x \notin (0, a) \\ 0 & x \in (0, a) \end{cases}$$

LST.16 | 5/8/09.4 | Abington | Wednesday |

we will consider the case with $R \ll a$.

The Schrödinger eq. in two dimensions is:

$$-\frac{\hbar^2}{2m} \left(\frac{\partial^2 \psi}{\partial x^2} + \frac{\partial^2 \psi}{\partial y^2} \right) + V(x,y) \psi(x,y) = E \psi(x,y).$$

using separation of variables.

$$\psi(x,y) = \phi(x) \phi(y)$$

$$\Rightarrow -\frac{\hbar^2}{2m} (\phi_{xx} \phi + \phi \phi_{yy}) = E \phi(x) \phi(y) \quad x \in (0,a)$$

$$\Rightarrow -\frac{\hbar^2}{2m} \left(\frac{\phi_{xx}}{\phi} + \frac{\phi_{yy}}{\phi} \right) = E = \text{const.}$$

$$\Rightarrow -\frac{\hbar^2}{2m} \frac{\phi_{xx}}{\phi} = C_1 \quad -\frac{\hbar^2}{2m} \frac{\phi_{yy}}{\phi} = C_2 \quad \text{with } C_1 + C_2 = E$$

we solve the two equations subject to the boundary conditions, with $\phi(y) = \phi(y + 2\pi R)$.

in x-direction we have as before.

$$\phi_1(x) = C_0 \sin \frac{\pi x}{a}$$

in y-direction | $\phi_2(y) = a_0 \sin \frac{l \cdot y}{R} + b_0 \cos \frac{l \cdot y}{R}$

for $l=0$ we have.

$$|\phi_0(y)| = b_0$$

LST, 17 | 5/8/09.5 | Abingdon / Wednesday |

Putting back the solutions into the equations for x and y

$$\Rightarrow E_{\pm l} = C_1 + C_2 = \frac{\hbar^2}{2m} \left[\frac{\hbar^2 \pi^2}{a^2} + \frac{l^2}{R^2} \right]$$

new: For $l \neq 0$ the spectrum is doubly degenerate $\pm l$.

For $l = 0$ the energy states are the same as the 1-dimensional case.

Lowest new state | The lowest new energy level is for $l = 1$ $l = 1$.

$$E_{1,1} = \frac{\hbar^2}{2m} \left[\left(\frac{\pi}{a} \right)^2 + \left(\frac{1}{R} \right)^2 \right]$$

in the case with $R \ll a$

$$E_{1,1} \sim \frac{\hbar^2}{2m} \frac{1}{R^2}$$

This energy is comparable to the l -level eigenstates of the one dimensional problem when

$$\frac{\hbar^2 \pi^2}{a} \sim \frac{1}{R} \rightarrow \hbar \sim \frac{1}{\pi} \frac{a}{R}$$

Since $R \ll a$ \hbar is very large.

\Rightarrow The first new energy levels are at an energy far above those of the low-lying original states.

\rightarrow The extra dimension remains hidden provided that $R \ll a$.

As we will see later strings introduce a new feature,

$$E \sim \frac{m^2}{R^2} + \frac{n^2 R^2}{\alpha'^2} \quad \textcircled{0}$$

strings can wrap around the compact dimension.

and create new states \sim wrapping modes,

For point particle \rightarrow no wrapping modes $E \sim \frac{m^2}{R^2}$,

Electromagnetism and gravitation in various dimensions

classical
electrodynamics

Maxwell's equations describe the dynamics of electric and magnetic fields in the presence of electric charges and currents.

$$\vec{\nabla} \times \vec{E} = -\frac{1}{c} \frac{\partial \vec{B}}{\partial t}$$

$$\vec{\nabla} \cdot \vec{B} = 0 \rightarrow \text{no magnetic charges}$$

$$\vec{\nabla} \cdot \vec{E} = \rho \rightarrow \text{electric field has charged sources}$$

$$\vec{\nabla} \times \vec{B} = \frac{1}{c} \vec{j} + \frac{1}{c} \frac{\partial \vec{E}}{\partial t} \rightarrow \text{moving charges create a magnetic field.}$$

\vec{E}, \vec{B} are the three dimensional electric and magnetic fields

ρ, \vec{j} are the electric charge and current densities.

Lorentz Law: $\frac{d\vec{P}}{dt} = q \left(\vec{E} + \frac{\vec{v}}{c} \times \vec{B} \right) \rightarrow$ effect of E & B fields on a moving charged particle.

LST, 19/5/18/09.7/Abingdon (Wednesday)

since $\vec{\nabla} \cdot \vec{B} = 0 \Rightarrow \vec{B} = \vec{\nabla} \times \vec{A}$

because
using our
summation
convention
and

$$\begin{aligned} (\vec{\nabla} \cdot (\vec{\nabla} \times \vec{A})) &= \partial_i \epsilon_{ijk} \partial_j A_k = \epsilon_{ijk} \partial_i \partial_j A_k = \\ &= -\epsilon_{jik} \partial_i \partial_j A_k = -\epsilon_{ijk} \partial_j \partial_i A_k = \\ &= -\epsilon_{ijk} \partial_i \partial_j A_k \Rightarrow \epsilon_{ijk} \partial_i \partial_j A_k = 0 \end{aligned}$$

$\epsilon_{ijk} = \begin{cases} +1 & \epsilon_{123} \\ -1 & \epsilon_{213} \\ 0 & \epsilon_{113} \\ \text{etc...} \end{cases}$
 $i, j, k = 1, 2, 3$

in electrostatics, $\vec{\nabla} \times \vec{E} = 0 \Rightarrow \vec{E} = -\vec{\nabla} \phi$

in electrodynamics $\vec{\nabla} \times \vec{E} = -\frac{1}{c} \frac{\partial}{\partial t} \vec{\nabla} \times \vec{A}$ (From Maxwell eq.)

$$\Rightarrow \vec{\nabla} \times \left(\vec{E} + \frac{1}{c} \frac{\partial \vec{A}}{\partial t} \right) = 0$$

$$\Rightarrow \vec{E} = -\vec{\nabla} \phi - \frac{1}{c} \frac{\partial \vec{A}}{\partial t}$$

\Rightarrow we can describe E & M dynamics in terms of the potential functions \vec{A} & ϕ .

Manifestly relativistic electrodynamics

$$A^\mu = (\underline{\phi}, A^1, A^2, A^3) \rightarrow 4\text{-vector}$$

$$A_\mu = (-\underline{\phi}, +\vec{A})$$

From A_μ we create the electromagnetic field strength

$$F_{\mu\nu} = (\partial_\mu A_\nu - \partial_\nu A_\mu)$$

$$\partial_\mu = \left(\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right) \quad \partial^\mu = \left(-\frac{1}{c} \frac{\partial}{\partial t}, \vec{\nabla} \right)$$

we have,

$$F_{\nu\mu} = \partial_\nu A_\mu - \partial_\mu A_\nu = -(\partial_\mu A_\nu - \partial_\nu A_\mu) = -F_{\mu\nu}$$

$\Rightarrow F_{\mu\nu}$ is antisymmetric.

\Rightarrow diagonal components of $F_{\mu\nu}$ vanish.

let i denote a spatial index i.e. $i = 1, 2, 3$.

Then $F_{0i} = \partial_t(A_i) - \partial_i A_0 = \frac{1}{c} \frac{\partial}{\partial t} (+A^i) + \partial \Phi = -E_i$

$$F_{ij} = \partial_i(A_j) - \partial_j(A_i) = +(\partial_i A_j - \partial_j A_i) = +(\nabla \times \vec{A})_k = +B_k$$

Hence

$$F_{\mu\nu} = \begin{pmatrix} 0 & -E_1 & -E_2 & -E_3 \\ +E_1 & 0 & +B_3 & -B_2 \\ +E_2 & -B_3 & 0 & +B_1 \\ +E_3 & +B_2 & -B_1 & 0 \end{pmatrix}$$

\rightarrow The electric and magnetic fields are encoded in

The field strength tensor $F_{\mu\nu}$

gauge transformations: The potential A^μ can be changed by gauge transformations

$$A_\mu \rightarrow A'_\mu = A_\mu + \partial_\mu \epsilon(x^\mu)$$

where $\epsilon(x^\mu)$ is a scalar function of x^μ .

without affecting $F_{\mu\nu}$.

$$F_{\mu\nu} \rightarrow F'_{\mu\nu} = \partial_\mu A'_\nu - \partial_\nu A'_\mu = \partial_\mu (A_\nu + \partial_\nu \epsilon) - \partial_\nu (A_\mu + \partial_\mu \epsilon) = \partial_\mu A_\nu - \partial_\nu A_\mu = F_{\mu\nu}$$

LST.21 | 5/8/09.9 | Abingdon | Wednesday

In terms of ϕ, \vec{A} we have.

$$\begin{aligned}\phi &\rightarrow \phi' = \phi - \frac{1}{c} \frac{\partial \epsilon}{\partial t} \\ \vec{A} &\rightarrow \vec{A}' = \vec{A} + \nabla \epsilon\end{aligned}$$

and \vec{E} and \vec{B} are unchanged.

$$\begin{aligned}\text{i.e. } \vec{E}' &= -\nabla \phi' - \frac{1}{c} \frac{\partial \vec{A}'}{\partial t} = -\nabla \left(\phi - \frac{1}{c} \frac{\partial \epsilon}{\partial t} \right) - \frac{1}{c} \frac{\partial (\vec{A} + \nabla \epsilon)}{\partial t} \\ &= -\nabla \phi - \frac{1}{c} \frac{\partial \vec{A}}{\partial t} + \frac{1}{c} \nabla \frac{\partial \epsilon}{\partial t} - \frac{1}{c} \frac{\partial \nabla \epsilon}{\partial t} = \vec{E}\end{aligned}$$

$$\text{and } \nabla \times \vec{A}' = \nabla \times (\vec{A} + \nabla \epsilon) = \nabla \times \vec{A} + \nabla \times \nabla \epsilon = \nabla \times \vec{A} = \vec{B}$$

$$\text{(because } \nabla \times \nabla \epsilon = \epsilon_{ijk} \partial_j \partial_k \epsilon = -\epsilon_{ikh} \partial_j \partial_k \epsilon = -\epsilon_{ikh} \partial_k \partial_j \epsilon = -\nabla \times \nabla \epsilon)$$

Next we write Maxwell's equations in a manifestly relativistic form.

consider $T_{\lambda\mu\nu} = \partial_\lambda F_{\mu\nu} + \partial_\nu F_{\lambda\mu} + \partial_\mu F_{\nu\lambda}$

$T_{\lambda\mu\nu}$ is identically zero

because $T_{\lambda\mu\nu} = \partial_\lambda (\cancel{\partial_\mu A_\nu} - \cancel{\partial_\nu A_\mu}) + \partial_\nu (\cancel{\partial_\lambda A_\mu} - \cancel{\partial_\mu A_\lambda}) + \partial_\mu (\cancel{\partial_\nu A_\lambda} - \cancel{\partial_\lambda A_\nu}) = 0$

hence $\partial_\lambda F_{\mu\nu} + \partial_\nu F_{\lambda\mu} + \partial_\mu F_{\nu\lambda} = 0$.

using $F_{\mu\nu} = -F_{\nu\mu}$ it follows that,

$$T_{\lambda\mu\nu} = -T_{\lambda\nu\mu} \quad T_{\lambda\mu\nu} = -T_{\mu\lambda\nu}$$

i.e. $T_{\lambda\mu\nu}$ is antisymmetric under exchange of any two indices.

LST.22 / 16/8/09.1 / Abingdon / Sunday

The source free Maxwell equations are obtained

from $T_{\lambda\mu} = \partial_\lambda F_{\mu\nu} + \partial_\mu F_{\nu\lambda} + \partial_\nu F_{\lambda\mu} \equiv 0$

Since $T_{\lambda\mu}$ is totally antisymmetric and $T_{\lambda\mu} \equiv 0$

\Rightarrow 4 independent equations.

e.g. $\lambda=0 \quad \mu=i \quad \nu=j$.

$$T_{0ij} = \partial_0 F_{ij} + \partial_i F_{j0} + \partial_j F_{0i} =$$

$$= +\partial_t(+B_k) + \partial_i(+E_j) - \partial_j E_i =$$

$$= +\partial_t B_k + \epsilon_{kij} \partial_i E_j = 0$$

$$\Rightarrow (\nabla \times \vec{E})_k + \frac{\partial (\vec{B})_k}{\partial t} = 0$$

FOR $\lambda=i \quad \mu=j \quad \nu=k$

$$T_{ijk} = \partial_i F_{jk} + \partial_j F_{ki} + \partial_k F_{ij}$$

take $i=1 \quad j=2 \quad k=3 \Rightarrow \partial_1 F_{23} + \partial_2 F_{31} + \partial_3 F_{12}$

$$= \partial_1(+B_1) + \partial_2(+B_2) + \partial_3(+B_3) = 0$$

$$\Rightarrow (\nabla \cdot \vec{B}) = 0$$

The two remaining Maxwell equations

contain electric charges and currents.

introduce: $j^\mu = (c\rho, j^1, j^2, j^3) \rightarrow$ FOUR vector current density.

LST. 23 | 16/R/09.2 | Abingdon | Saturday |

$\rho \rightarrow$ charge density.

$\vec{J} = (j_1, j_2, j_3)$ current density.

introduce $F^{\mu\nu} = \eta^{\mu\alpha} \eta^{\nu\beta} F_{\alpha\beta}$

Show (1) $F^{\mu\nu} = -F^{\nu\mu}$ $F_{0i} = -F_{i0}$ $F^{ij} = F_{ij}$.

$$2) F^{\mu\nu} = (\partial^\mu A^\nu - \partial^\nu A^\mu)$$

$$3) F^{\mu\nu} = \begin{pmatrix} 0 & +E_1 & +E_2 & +E_3 \\ -E_1 & 0 & +B_3 & -B_2 \\ -E_2 & -B_3 & 0 & +B_1 \\ -E_3 & +B_2 & -B_1 & 0 \end{pmatrix}$$

The remaining Maxwell equations then take the form.

$$\partial_\nu F^{\mu\nu} = \partial^\nu F^{\mu\nu} = \frac{1}{c} j^\mu \quad (\text{HW. 2})$$

In the absence of sources.

$$\Rightarrow \partial_\nu F^{\mu\nu} = 0 \rightarrow \partial_\nu \partial^\mu A^\nu - \partial_\nu \partial^\nu A^\mu = 0.$$

$$\text{OR } \partial^\nu \partial^\mu A^\nu - \partial^\mu (\partial^\nu A^\nu) = 0.$$

where $\partial_\nu \partial^\nu \equiv \partial^2$.

we can now use these equations to generalise.

Maxwell equations to any dimensions.

4D $0, 1, 2, 3.$ $\eta^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & +1 \end{pmatrix}$

nD $0, 1, 2, \dots, n$ $\eta^{\mu\nu} = \begin{pmatrix} -1 & & & \\ & +1 & & \\ & & +1 & \\ & & & \dots \\ & & & & +1 \end{pmatrix}_{n \times n}$

Example: 3D space-time $\rightarrow 1+2 \rightarrow \begin{pmatrix} -1 \\ +1 \\ +1 \end{pmatrix}$

$$F_{\mu\nu} = \begin{pmatrix} 0 & -E_x & -E_y \\ +E_x & 0 & +B_z \\ +E_y & -B_z & 0 \end{pmatrix}$$

in D dimensions: $F^{0i} = -F_{0i} = +E_i \rightarrow$ electric field
 $F_{ij} = -F_{ji} \Rightarrow$ magnetic field

in D-dimension E_i is a spatial vector $i=1, \dots, D$
 B_i is not a spatial vector.

spheres in higher dimensions

a 2-sphere in 3 Dimensions defined by the equation,

$$S^2(R) = x_1^2 + x_2^2 + x_3^2 = R^2$$

The volume enclosed by the sphere is a ball.

$$B^3(R) = x_1^2 + x_2^2 + x_3^2 \leq R^2$$

where R is the radius of the ball.

similarly in D-dimensions.

$$S^{D-1}(R) = x_1^2 + x_2^2 + \dots + x_D^2 = R^2$$

$$B^D(R) = x_1^2 + x_2^2 + \dots + x_D^2 \leq R^2$$

$$\sqrt{x_1^2 + x_2^2} = R^2$$

the volume of a 1-sphere: $\text{vol}(S^1(R)) = 2\pi R$ (length)

" " " a 2-sphere: $\text{vol}(S^2(R)) = 4\pi R^2$ (area)

" " " a D-1 sphere: $\text{vol}(S^{D-1}(R)) = \text{vol}(S^{D-1}) R^{D-1}$