Holography for non-relativistic CFTs Herzog, Rangamani & SFR, 0807.1099, Rangamani, Son, Thompson & SFR, 0811.2049, SFR & Saremi, 0907.1846

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#### Outline

- Review holography for relativistic theories
- Application to condensed matter physics
- Holography for non-relativistic theories: Lifshitz example
- Action principle
- Stress tensor
- Schrödinger example
- Action & stress tensor for Schrödinger

#### Holographic description of CFTs

• Identify SO(d, 2) conformal symmetry with isometries of  $AdS_{d+1}$ . In Poincare coordinates.

$$ds^2 = r^2 \eta_{\mu\nu} dx^{\mu} dx^{\nu} + \frac{dr^2}{r^2},$$

Dilatation is  $D: x^{\mu} \to \lambda x^{\mu}, r \to \lambda^{-1}r$ .

• 
$$\langle e^{\int \phi_0 \mathcal{O}} \rangle = Z_{string}[\phi_0] \approx e^{-S[\phi_0]}.$$

- Classical gravity valid at strong 't Hooft coupling.
- Many explicit examples:  $AdS_5 \times X^5$ .
- Focus on universal subsector: pure gravity in bulk. Boundary stress tensor  $\langle T_{\mu\nu} \rangle = \frac{1}{\sqrt{\hbar}} \frac{\delta S}{\delta \hat{h}^{\mu\nu}}$  Henningson Skenderis Balasubramanian
- Finite temperature described by bulk black hole solution.

Kraus

Application to condensed matter physics

#### Application to QCD long studied. Why condensed matter?

- Rich system:
  - CFTs arise as IR desc near critical points; often strongly coupled.
  - Many different theories; can tune Hamiltonian in some settings.
- AdS/CFT provides a useful new perspective:
  - Few other methods for calculation at strong coupling.
  - Gravity dual calculates observables like transport coefficients directly, no quasiparticle picture.
- Prompts new questions:
  - Charge transport, phase transitions
  - Can have theories with an anisotropic scaling symmetry D:  $x^i \rightarrow \lambda x^i, t \rightarrow \lambda^z t$ .

## Application to condensed matter physics

#### However, this is a big extension of AdS/CFT.

- Known examples involve large N gauge theories. To apply these ideas to condensed matter, need to assume other field theories will have holographic duals.
- Proceed phenomenologically: construct bulk theories with qualitative features of interest.
- Big question: for what types of theories can we construct a bulk dual in the classical approximation?
  - Strong coupling
  - $N_{dof} \gg 1$
  - Hierarchy in spectrum:  $\Delta_{s>2} \gg \Delta_{s\leq 2}$ .

## Non-relativistic holography

In condensed matter, have theories with an anisotropic scaling symmetry D:  $x^i \rightarrow \lambda x^i, t \rightarrow \lambda^z t$ .

Two cases of interest:

- Lifshitz-like theories:  $D, H, \vec{P}, M_{ij}$
- Schrödinger symmetry: add Galilean boosts  $\vec{K}$ . z = 2 special.

Can these also have a dual geometrical description?

## Non-relativistic holography

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Lifshitz:

• Simple deformation of AdS:

Kachru Liu Mulligan

$$ds^{2} = -r^{2z}dt^{2} + r^{2}d\mathbf{x}^{2} + \frac{dr^{2}}{r^{2}}.$$

#### Lifshitz geometry

• Solution of a theory with a massive vector:

$$S=\int d^4x\sqrt{-g}(R-2\Lambda-rac{1}{4}F_{\mu
u}F^{\mu
u}-rac{1}{2}m^2A_\mu A^\mu),$$

with  $\Lambda = -\frac{1}{2}(z^2 + z + 4)$ ,  $m^2 = 2z$ . Lifshitz solution has

$$A = \alpha r^z dt, \quad \alpha^2 = \frac{2(z-1)}{z}.$$

• Finite temperature black hole solutions obtained numerically.



 Analytic black holes in a theory with an extra scalar field. Balasubramanian McGreevy

• Embedded in string theory yesterday!

Hartnoll Polchinski Silverstein Tong

## Asymptotically Lifshitz spacetimes

Need to specify boundary conditions for metric,  $A_{\mu}$ .

- Different metric components scale differently with *r*.
   ⇒ No conformal boundary.
- Use frame fields. For Lifshitz geometry, natural basis is

$$e^{(0)} = r^z dt, \quad e^{(i)} = r dx^i, \quad e^{(3)} = \frac{dr}{r},$$

• Write gauge field as

 $A = \alpha e^{(0)}.$ 

#### Asymptotically Lifshitz spacetimes

At large r, assume spacetime is Lifshitz + linearized corrections. By choice of frame, write

$$e^{(0)} = r^{z} \hat{e}^{(0)} = r^{z} [(1 + \frac{1}{2} \hat{h}_{tt})dt + v_{1i}dx^{i}],$$

$$e^{(i)} = r \hat{e}^{(i)} = r [v_{2i}dt + (\delta^{i}_{\ j} + \frac{1}{2} \hat{h}^{i}_{\ j})dx^{j}], \quad e^{(3)} = \frac{dr}{r},$$

$$A^{M} = \alpha (1 + \hat{a}_{t})\delta^{M}_{0} + \alpha \hat{a}_{r}\delta^{M}_{3}.$$
y conditions  $\hat{h}_{tt}, v_{1i}, v_{2i}, \hat{h}_{ij}, \hat{a}_{t} \to 0$  as  $r \to \infty$ .

Boundary conditions  $h_{tt}$ ,  $v_{1i}$ ,  $v_{2i}$ ,  $h_{ij}$ ,  $\hat{a}_t \rightarrow 0$  as  $r \rightarrow \infty$ . For  $1 \le z < 2$ , also impose  $r^{\frac{1}{2}(z+2-\beta_z)}\hat{a}_t \rightarrow 0$  as  $r \rightarrow \infty$ , where  $\beta_z^2 = (z+2)^2 + 8(z-1)(z-2)$ .

#### Linearized solutions

| Bertoldi   | SFR<br>Saremi |  |
|------------|---------------|--|
| Burrington |               |  |
| Peet       | Sarenn        |  |

Constant modes

$$\begin{aligned} \hat{a}_t, \hat{h}_{tt}, \hat{h}_i^i &\sim \frac{a_1}{r^{z+2}}, \ \frac{a_2}{r^{\frac{1}{2}(z+2+\beta_z)}} \\ v_{1i} &= \frac{c_{2i}}{r^{z+2}} + \frac{c_{3i}}{r^{3z}}, \quad v_{2i} &= \frac{(z^2 - 4)}{z(z-4)}c_{2i}r^{z-4} + \frac{3z}{(z+2)}\frac{c_{3i}}{r^{z+2}}, \\ \hat{h}_{ij}^T &= \frac{t_{ij}}{r^{z+2}}. \end{aligned}$$

(Non-constant modes a little more complicated.)

Expect a normalizable mode for each field, giving expectation value of corresponding operator.

- Scalars ok; scale inv implies one relation, expect 2 modes.
- For vectors,  $c_{2i}$  violates boundary condition on  $v_{2i}$  for z > 4. Fewer normalizable modes than expected??

Action

#### SFR Saremi

AdS/CFT dictionary is based on an action:  $\langle e^{\int \phi_0 \mathcal{O}} \rangle \approx e^{-S_{on-shell}[\phi_0]}$ . Need an action *S* such that for asymptotically Lifshitz spacetimes

• S finite on-shell.

•  $\delta S = 0$  for arbitrary variations preserving the boundary conditions. Naive bulk action divergent. Regularize by adding boundary terms:

$$S_{ct} = bulk + rac{1}{16\pi G_4} \int d^3\xi \sqrt{-h} (2K - 4 - \sqrt{2z(z-1)}\sqrt{-A_{\alpha}A^{\alpha}})$$

 $\triangleright$  Need derivative terms for asymptotically Lifshitz spacetimes: consider

$$S_{deriv} = \int d^3 \xi \sqrt{-h} [\sigma_1 R^h + \sigma_2 \nabla_\alpha A_\beta \nabla^\alpha A^\beta + \sigma_3 (\Box A_\alpha) (\Box A^\alpha)].$$

 $\star~S=S_{ct}+S_{deriv}$  satisfies  $\delta S=0$  for variations satisfying our boundary conditions,  $S_{on-shell}$  finite.

#### Stress tensor

In relativistic case, stress tensor defined by  $\delta S = \int d^3\xi T_{\alpha\beta}\delta \hat{h}^{\alpha\beta}$ . Two choices for vector field:

- Holding  $A_{\alpha}$  fixed or holding  $A_A$  fixed.
- Latter gives conserved charges gen bdy diffeos.

Hollands Ishibashi Marolf

• In non-relativistic case, it is latter that gives finite results.

Non-relativistic theory: energy and momentum not related by a coordinate transformation. Stress-energy complex

- Energy density  $\mathcal{E}$ , energy flux  $\mathcal{E}_i$
- Momentum density  $\mathcal{P}_i$ , stress tensor  $\Pi_{ij}$ .

Conservation equations  $\partial_t \mathcal{E} + \partial_i \mathcal{E}^i = 0$ ,  $\partial_t \dot{\mathcal{P}}_j + \partial_i \Pi^i{}_j = 0$ . Define by

$$\delta S = \int [-\mathcal{E}\delta \hat{e}_t^{(0)} - \mathcal{E}^i \delta \hat{e}_i^{(0)} + \mathcal{P}_i \delta \hat{e}_t^{(i)} + \Pi_i^j \delta \hat{e}_j^{(i)} + s_A \delta A^A].$$

Results finite: *E* ∝ *a*<sub>1</sub>, *E<sub>i</sub>* ∝ *c*<sub>3i</sub>, *P<sub>i</sub>* ∝ *c*<sub>2i</sub>, Π<sub>ij</sub> ∝ *a*<sub>1</sub>δ<sub>ij</sub> + *t<sub>ij</sub>*.
Bulk equations of motion imply conservation equations, *zE* = Π<sup>i</sup><sub>i</sub>.

# Schrödinger symmetry

Galilean symmetry: rotations  $M_{ij}$ , translations  $P_i$ , boosts  $K_i$ , (i = 1, ..., d), Hamiltonian H, particle number N. Extended by the dilatation D,

 $[D, P_i] = iP_i, [D, H] = ziH, [D, K_i] = (1 - z)iK_i, [D, N] = i(2 - z)N.$ 

Dynamical exponent z determines scaling of H under dilatations.

For z = 2, N is central, and there is a special conformal generator C: [D, C] = -2iC, [H, C] = -iD.

Schrödinger algebra: Symmetry of free Schrödinger equation.

\* Isometries of a gravitational dual?

#### Geometrical dual

Embed Galilean symmetry in ISO(d + 1, 1) by light-cone quant:  $H = \tilde{P}_+$ ,  $N = \tilde{P}_-$ ,  $K_i = \tilde{M}_{-i}$ . Extend to embed Sch(d) in SO(d + 2, 2) by

$$D = ilde{D} + (z-1) ilde{M}_{+-}$$

\* Sch(d) is a subgroup of SO(d+2,2)

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Gravitational dual: deform  $AdS_{d+3}$  to

$$ds^{2} = -r^{4}(dx^{+})^{2} + r^{2}(-2dx^{+}dx^{-} + d\mathbf{x}^{2}) + \frac{dr^{2}}{r^{2}}.$$

• Solution of a theory with a massive vector,  $A_+ = r^2$ .

• *N* discrete implies *x*<sup>-</sup> periodic. Compact null direction?

#### Dual in string theory

| Herzog    | Adams           | Maldacena |
|-----------|-----------------|-----------|
| Rangamani | Balasubramanian | Martelli  |
| SFR       | McGreevy        | Tachikawa |

Take  $AdS_5 \times S^5$ ,

$$ds^{2} = r^{2}(-2dx^{+}dx^{-} + d\mathbf{x}^{2}) + \frac{dr^{2}}{r^{2}} + (d\psi + P)^{2} + d\Sigma_{4}^{2},$$

and apply a TsT transformation:

• T-dualize the Hopf fiber coordinate  $\psi$  to  $\tilde{\psi}$ ,

• Shift 
$$x^- o ilde{x}^- = x^- + ilde{\psi}$$
,

• T-dualise  $\tilde{\psi}$  to  $\psi$  at fixed  $\tilde{x}^-$ .

Resulting solution is

$$ds^{2} = -r^{4}(dx^{+})^{2} + r^{2}(-2dx^{+}dx^{-} + d\mathbf{x}^{2}) + \frac{dr^{2}}{r^{2}} + (d\psi + P)^{2} + d\Sigma_{4}^{2},$$

 $B=r^2dx^+\wedge (d\psi+P).$ 

Maldacena Martelli Tachikawa

#### Heating up Schrödinger

Apply the TsT transformation to the Schwarzschild-AdS solution,

$$ds^{2} = r^{2}(-f(r)dt^{2} + dy^{2} + dx^{2}) + \frac{dr^{2}}{r^{2}f(r)} + ds^{2}_{5^{5}},$$

where  $f(r) = 1 - r_+^4/r^4$ . Resulting solution in 5d is

$$ds^{2} = r^{2}k(r)^{-\frac{2}{3}} \left( \left[ \frac{r_{+}^{4}}{4\beta^{2}r^{4}} - r^{2}f(r) \right] (dx^{+})^{2} + \frac{\beta^{2}r_{+}^{4}}{r^{4}} (dx^{-})^{2} - [1 + f(r)]dx^{+}dx^{-} \right) + k(r)^{\frac{1}{3}} \left( r^{2}dx^{2} + \frac{dr^{2}}{r^{2}f(r)} \right).$$

$$A = \frac{r^{2}}{k(r)} \left( \frac{1 + f(r)}{2} dx^{+} - \frac{\beta^{2}r_{+}^{4}}{r^{4}} dx^{-} \right),$$

$$e^{\phi} = \frac{1}{\sqrt{k(r)}}; \quad k(r) = 1 + \frac{\beta^{2}r_{+}^{4}}{r^{2}}.$$

 $\triangleright \beta$  param. choice of  $x^-$  coord in TsT; define  $\gamma^2 \equiv \beta^2 r_+^4$ .

## Action principle

| Herzog           | SFR<br>Saremi |
|------------------|---------------|
| Rangamani<br>SFR |               |
| SER              |               |

Similarly construct action by adding boundary counterterms:

 $S = bulk + \int d^4 \xi \sqrt{-h} (2K - 6 + (1 + c\phi)A^{\mu}A_{\mu} + (2c + 3)\phi^2)$ 

Finite, satisfies  $\delta S = 0$ , but boundary conditions for variations more restrictive.

Stress tensor — proceed as in Lifshitz case: vary metric at fixed  $A^A$ ,  $\phi$ . Differences:

- Conserved particle number. Particle number density  $\rho$ , flux  $\rho_i$ .
- No natural choice of frame.

Write

$$\delta S = \int d^4 \xi (s_{\alpha\beta} \delta h^{\alpha\beta} + s_\alpha \delta A^\alpha),$$

Define  $\mathcal{E} = 2s_{+}^{+} - s_{-}^{+}A_{+}, \quad \mathcal{E}^{i} = 2s_{+}^{i} - s_{-}^{i}A_{+}, \text{ etc.}$ 

- For black hole solution,  $\mathcal{E} = r_+^4$ ,  $\Pi_{xx} = \Pi_{yy} = r_+^4$ ,  $\rho = 2\beta^2 r_+^4$ .
- For solutions obtained by TsT from vacuum AdS solution, agrees with AdS stress tensor.

Simon Ross (Durham)

#### Discussion

- NRCFT is an interesting and challenging extension of AdS/CFT.
- Action and stress tensor constructed: first steps in holographic dictionary.

Open questions:

- Characterising field theories that have a weakly curved gravity dual.
- One-point function for massive vector.
- Role of extra  $x^-$  direction in Schrödinger geometry.
- Condensed matter applications: relation to strange metals in paper yesterday.