Calculating the Hadronic Vacuum Polarisation Contribution to the Anomalous Magnetic Moment of the Muon,  $a_{\mu}^{had,VP}$ 

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### Overview

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### The Standard Model (SM): An Incomplete Theory

- $a_{\mu}$  is one of the most precisely measured quantities in particle physics, accurate to 0.54ppm.
- $a_{\mu}^{SM} < a_{\mu}^{exp}$  by approximately 3 standard deviations.
- Any deviation could herald the existence of as-yet-unknown new physics beyond the SM.
- Experiment will be four times more accurate after the completion of Muon g-2 experiment at Fermilab!



By improving the precision of  $a_{\mu}$ , we can determine whether the SM is an incomplete theory!!

#### Motivation

### The Anomalous Magnetic Moment

• A charged elementary particle generates an angular momentum,  $\vec{\mu}$ 

$$\vec{\mu} = g \frac{e}{2m} \vec{s}$$

- In 1928, Dirac predicted that g = 2 for spin-1/2 particles.
- However, due to radiative corrections

$$a_l = rac{(g-2)}{2}$$
 where  $(l=e,\mu, au)$ 

 In 1948, Schwinger calculated the leading order (LO) contribution as  $a_l^{QED,LO} = \frac{\alpha}{2\pi}$  .

### Contributions from the SM

$$a_{\mu} = \frac{(g-2)}{2} = a_{\mu}^{QED} + a_{\mu}^{EW} + a_{\mu}^{had} + a_{\mu}^{NewPhysics?}$$

$$\begin{split} a^{QED}_{\mu} &= (11658471.808 \pm 0.015) \times 10^{-10} \quad \text{[Kinoshita et al.]} \\ a^{EW}_{\mu} &= (15.4 \pm 0.2) \times 10^{-10} \quad \text{[Czarnecki et al., Knecht et al.]} \\ a^{had}_{\mu} &= (695.6 \pm 4.9) \times 10^{-10} \quad \text{[HLMNT 2012]} \\ a^{SM}_{\mu} &= a^{QED}_{\mu} + a^{EW}_{\mu} + a^{had}_{\mu} = (11659182.8 \pm 4.9) \times 10^{-10} \\ a^{exp}_{\mu} &= (11659208.9 \pm 6.3) \times 10^{-10} \quad \text{[arXiv:1001.2898]} \\ &\therefore \mathbf{a}^{exp}_{\mu} - \mathbf{a}^{SM}_{\mu} = (\mathbf{26.1 \pm 8.0}) \times \mathbf{10^{-10}} \end{split}$$

corresponding to a  $3.3\sigma$  discrepancy.

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Calculating  $a_{\mu}^{had,VP}$ 

### The Hadronic Contribution

Uncertainties from the hadronic sector completely dominate  $\Delta a_{\mu}^{SM}$ !



Of these,  $a_{\mu}^{\rm had,\ VP\ LO}$  has the largest uncertainty.

The internal hadronic "blob" contains the contributions from all the possible hadronic states!!

## Calculating $a^{had}_{\mu}$

• By means of unitarity,

$$\sigma_{e^+e^- \to hadrons}(s) = \left(\frac{4\pi\alpha}{s}\right) {\rm Im} \Pi^{\gamma}_{had}(s)$$



• Using analyticity, we get the dispersion integral

$$a_{\mu} = \frac{\alpha}{\pi} \int_0^\infty \frac{ds}{s} \frac{1}{\pi} \mathrm{Im} \Pi_{had}^{\gamma}(s) K(s)$$

and 
$$a_{\mu}^{had,LOVP} = \left(\frac{\alpha m_{\mu}}{3\pi}\right)^2 \int_{s^{th}}^{\infty} ds \; \frac{R_{had}^{data}(s)K(s)}{s^2}$$

where  $R_{had}(s) \equiv \sigma^0_{e^+e^- \to hadrons} / \frac{4\pi \alpha^2}{3s}$  and  $K(s) = \frac{m_{\mu}^2}{3s} (0.4 \dots 1).$ 

• Integration limits go from  $s=4m_{\pi}^2\rightarrow\infty$  as  $K(s)=0.4\rightarrow1.$ 

### Combining Data Sets

#### Large number of available data sets for a given hadronic channel.

```
Z 75 !KLOE10
Z 337 !BaBar, PRL103 231801 (2009); 337pts, 0.30 .. 3.0
Z 60 !KLOE08, 0809.3950v2 (2008); 0.5916 .. 0.9747
v 10 !CMD-2. Akhmetshin06. hep-ex/0610016: 10 pts. 0.37 .. 0.52
y 29 !CMD-2, Akhmetshin06, hep-ex/0610021; 29 pts, 0.6 .. 0.97
y 45 ISND, Achasov04, re-analysis, hep-ex/0605013; 45 pts, 0.39 .. 0.97
y 36 !CMD-2, Aulchenko06, hep-ex/0603021, JETP Lett82(2005)743, 36pts, 1.04-1.38
n 60 !KLOE05, Rad.Ret., KLOE Note n.192, 60 pts, 0.35 .. 0.94 GeV^2 (OLD KLOE Data, exclude)
y 2 !OLYA-VEPP2, Vasserman79, SJNP30(1979)519; 2 pts, 0.4 and 0.44
y 4 !TOF-VEPP 2M, Vasserman81, SJNP33(1981)368; 4 pts, 0.4 .. 0.46
v 4 !NA7-CERN. Amendolia84. PL138B(1984)454: 4 pts. 0.32 .. 0.422
v 77 !OLYA-VEPP 2M. Barkov85. NPB256(1985)365: 77 pts. 0.6426 .. 1.397
v 24 !CMD-VEPP 2M. Barkov85, NPB256(1985)365; 24 pts. 0.36 .. 0.82
y 16 !DM1-ACO, Quenzer78, PL76B(1978)512; 16 pts, 0.483 .. 1.096
y 17 !DM2-DCI, Bisello89, PLB220(1989)321; 17 pts, 1.35 .. 2.125
y 13 !BCF-ADONE, Bollini75, Lett.NC14(1975)418; 13 pts, 1.2 .. 3.
y 1 !MEA-ADONE-Frascati, Esposito77, PLB67(1977)239; 1pt at 1.6
v 1 !MEA-ADONE-Frascati, Esposito80, NCL28(1980)337; 1pt at 1.45..1.52
v 8 !BCF-ADONE. Bernardini73. PL46B(1973)261: 8 pts. 1.2 .. 3.
v 28 !NOVOSIBIRSK-CMD, Dolinskv91, PRep202(1991)99; 28 pts. 0.81 .. 1.39
y 29 !NOVOSIBIRSK-VEPP-2M, OLYA, Bukin78, PL73B(1978)226; 29 pts, 1.33 .. 0.78
y 17 !NOVOSIBIRSK-VEPP-2M, OLYA, Koop79, INP-79-67 (1979); 17 pts, 1.06 .. 1.40
y 3 !Cosme85, LAL-1287 (1985) and NPB256(1985)365; 3 pts, 0.915, 0.99, 1.076 !should be from 1976
y 43 !CMD-2, REANALYSIS, hep-ex/0308008; 43 pts, 0.6105 .. 0.96152
y 60 !KLOE, Babusci 2013, PL 720B,336; 60 pts, 0.5958 .. 0.9721
```

#### We must make sure that for each data set we,

- correctly remove the photon VP effects.
- calculate accurate errors considering our VP corrections
- account for any final state radiative (FSR) corrections.

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How do we use data from different experiments in a given channel to compute its contribution to  $a_{\mu}^{had}$ ?

- Combine the data in a given channel before integrating.
- Re-bin the data points into *energy clusters*.

$$R_m = \left[\sum_{k} \sum_{i=1}^{N^{(k,m)}} \frac{R_i^{(k,m)}}{(d\tilde{R}_i^{(k,m)})^2}\right] \left[\sum_{k} \sum_{i=1}^{N^{(k,m)}} \frac{1}{(d\tilde{R}_i^{(k,m)})^2}\right]^{-1}$$
$$E_m = \left[\sum_{k} \sum_{i=1}^{N^{(k,m)}} \frac{E_i^{(k,m)}}{(d\tilde{R}_i^{(k,m)})^2}\right] \left[\sum_{k} \sum_{i=1}^{N^{(k,m)}} \frac{1}{(d\tilde{R}_i^{(k,m)})^2}\right]^{-1}$$

 $N^{(k,m)} = \text{total no. of data points within cluster } m.$ 

- Use adaptive clustering algorithm to produce *target clusters*.
  - $\rightarrow$  too small a cluster = precise data overwhelmed
  - $\rightarrow$  too large a cluster = data missed about resonance peaks

### Fitting The Data

Use a non-linear non-linear  $\chi^2$ -function [HLMNT, 2012]

$$\begin{split} \chi^2(R_m, f_k) &= \sum_{k=1}^{N_{exp}} \left( \frac{1 - f_k}{df_k} \right)^2 + \left\{ \sum_{m=1}^{N_{clu}} \sum_{i=1}^{N^{(k,m)}} \left( \frac{R_i^{(k,m)} - f_k R_m}{d\tilde{R}_i^{(k,m)}} \right)^2 \right\}_{\text{w/o cov. mat}} \\ &+ \left\{ \sum_{m=1}^{N_{clu}} \sum_{i=1}^{N^{(k,m)}} \sum_{j=1}^{N^{(k,m)}} \left( R_i^{(k,m)} - f_k R_m \right) C^{-1}(m_i, n_j) \left( R_j^{(k,n)} - f_k R_n \right) \right\} \end{split}$$

where  $N_{clu}$ ,  $N_{exp}$  are total number of clusters, experiments.

- Treat the statistical/systematic errors according to experimental data.
- Input covariance matrices where provided (last term!).

### Goodness-of-Fit

Minimise  $\chi^2(R_m, f_k)$  to give

- Fitted values  $\bar{R}_m$  and  $\bar{f}_k$ .
- $\bullet$  Minimised global  $\chi^2_{\rm min}$
- $\bullet\,$  Covariance matrix V(m,n) that defines the correlation of errors between the clusters.

$$V(m,n) = (dR_m)(dR_n)\rho_{\rm corr}(m,n)$$

 $\rightarrow$  Determine a 'goodness' of our overall fit from

$$\frac{\chi^2_{\rm min}}{{\rm d.o.f.}} = \frac{\chi^2_{\rm min}}{N_{\rm tot} - N_{\rm clu} - N_{\rm exp}}$$

# Local $\chi^2_m$ [HLMNT, 2012]



Channels	Global $\chi^2/dof$	Global inf err	Local inf err	Global - Local
$2\pi$	1.3971	3.0560	3.0896	-0.0336
$3\pi$	2.5669	1.1865	0.8045	+0.3820
$4\pi(2n)$	1.2922	1.1915	1.2593	-0.0678
$4\pi$	1.6926	0.4902	0.4738	+0.0164
$K^+K^-$	1.8423	0.5910	0.4781	+0.1129
$K_{S}^{0}K_{L}^{0}$	0.8055	0.1602	0.1627	-0.0025
$6\pi(2n)$	4.0304	0.3945	0.2383	+0.1562
$5\pi(1n)$	1.3831	0.0903	0.0857	+0.0046
$kk\pi\pi$	N/A	1.3234	1.3225	+0.0009

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### Using the Trapezoid Rule

Between two arbitrary energies  $E_a$  and  $E_b$ 

$$I = \int_{E_b^2}^{E_a^2} \frac{ds}{s} \ R_{had}(s) K(s) = 2 \int_{E_b^2}^{E_a^2} \frac{dE}{E^2} \ E \ R_{had}(E^2) K(E^2) = \bar{I} \qquad \text{with error} \pm \Delta \bar{I}$$

Then, suppose that  $E_m < E_a < E_{m+1}$  which is less than  $E_{n-1} < E_b < E_n$  and estimate using linear interpolation,

$$\bar{I} = 2\left(\frac{E_{m+1} - E_a}{2E_a}\bar{R}_aK_a + \frac{E_{m+2} - E_a}{2E_{m+1}}\bar{R}_{m+1}K_{m+1}\right) + 2\left(\sum_{l=m+2}^{n-2}\frac{E_{l+1} - E_{l-1}}{2E_l}\bar{R}_lK_l\right) + 2\left(\frac{E_b - E_{n-2}}{2E_{n-1}}\bar{R}_{n-1}K_{n-1} + \frac{E_b - E_{n-1}}{2E_b}\bar{R}_bK_b\right)$$

• We find the integral error  $\Delta \overline{I}$  via the inflated covariance matrix  $\tilde{V}$ .

$$(\Delta \bar{I})^2 = \sum_{p=m}^n \sum_{q=m}^n \frac{\partial \bar{I}}{\partial \bar{R}_p} \tilde{V}(p,q) \frac{\partial \bar{I}}{\partial \bar{R}_q}$$

where we must be careful that the border terms are well defined.

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 $e^+e^- 
ightarrow \pi^+\pi^-$ , [HLMNT, 2012]

 $\rightarrow$  Most important channel (> 70%)



### $e^+e^- ightarrow K^+K^-$ , [PhysRevD.88.032013]

- 159 New Points (14 sets, 364 points in total)
   → Improved precision in tails and around resonance.
- Systematic covariance matrix required.
- First analysis indicates definite reconsideration of data treatment to improve fit.



Figure : BABAR plot vs. New/Old fit

### Conclusions

- Improving accuracy of  $a_{\mu}$  provides stringent test of the SM and the possibly of new physics.
- Hadronic leading-order VP sector provides largest uncertainty.
- Using unitarity and analyticity, we can calculate  $a_{\mu}^{had}$  via a dispersion relation.
- Combine data sets through locally inflated target clusters.
- Employ a  $\chi^2$ -function to minimise and fit data.
- Once integrated, we can produce a new value for  $a_{\mu}^{had}$  and compare.
- Imminent emergence of new data promises for exciting years to come...

# Thank You