

# Magnetic moment $(g - 2)_\mu$ and SUSY

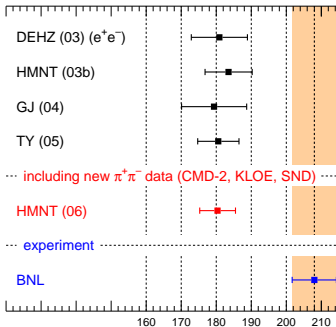
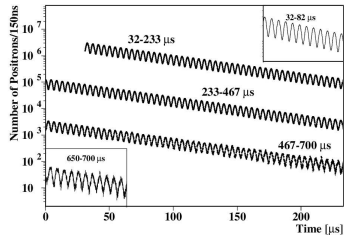
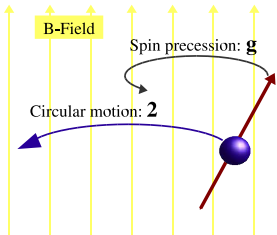
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SUPA, Edinburgh

BSMUK, Liverpool, March 2007

[DS '07]

# $(g - 2)$ : Magnetic Moment of the Muon



$$a_{\mu}(\text{exp}) = 11\,659\,208(6) \times 10^{-10}$$

$$a_{\mu}(\text{exp} - \text{SM}) = 28(8) \times 10^{-10}$$

**3.4 $\sigma$  deviation from SM-prediction!**

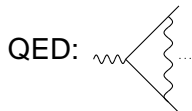
# Outline

- 1 Status: SM and experiment
- 2  $a_\mu$  and SUSY
- 3 Numerical results in SUSY
- 4 Campaign for new, better measurement
- 5 Conclusions

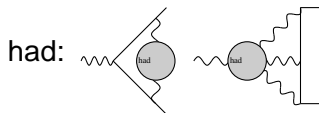
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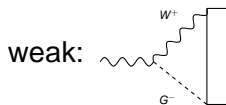
# Classification of SM contributions



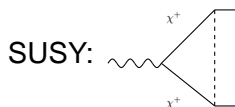
$$\sim 10^{-3} \quad 7000\sigma$$



$$\sim 10^{-7} \quad 100\sigma$$



$$\sim 10^{-9} \quad 2.5\sigma$$



$$\sim 10^{-8} \dots 10^{-10} \quad O(10\sigma)$$

## Era of the Brookhaven experiment

'78 – 2001: Theory precision increased to  $\pm 10 \times 10^{-10}$

2001: BNL experiment:  $a_{\mu}^{\text{exp}} - a_{\mu}^{\text{th}} = 43 \pm 16 \quad 2.7\sigma$

2002: correction of |b| sign error:  $28 \pm 16 \quad 1.7\sigma$

2002: BNL experiment:  $26 \pm 11 \quad 2.4\sigma$

2003: 4-Loop error, new LBL result:  $20 \pm 11 \quad 1.8\sigma$

'03–'06: Consolidation of SM result by effort of many groups

spring 2006: BNL experiment (final):  $24 \pm 10 \quad 2.4\sigma$

>summer 2006: new SM evaluations: **spectacular improvement**

# Current status

- Exp: finalized

- Th:

- new SM evaluations, based on new exp data for  $a_{\mu}^{\text{had}}$ :

$$a_{\mu}(\text{Exp-SM}) = \left\{ \begin{array}{ll} [\text{HMNT06}] & 28(8) \\ [\text{DEHZ06}] & 28(8) \\ [\text{FJ07}] & 29(9) \end{array} \right\} \times 10^{-10}$$

- better agreement between evaluations, **more precise**,  
**larger deviation from exp than ever before**

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# $a_\mu$ and SUSY

Two questions:

Could SUSY be the origin of the  $(28 \pm 8) \times 10^{-10}$  deviation?

Which restrictions on SUSY follow from (e.g.  $3\sigma$  band)

$$3 \times 10^{-10} < a_\mu^{\text{SUSY}} < 51 \times 10^{-10}?$$

# $g - 2$ in the MSSM

Key to understand  $g - 2$ :

## chiral symmetry

$g - 2 =$  chirality-flipping interaction

$$\bar{u}_R(p') \frac{\sigma_{\mu\nu} q^\nu}{2m_\mu} u_L(p) + (L \leftrightarrow R)$$

in each Feynman diagram we need to pick up one transition

$$\mu_L \rightarrow \mu_R \text{ or } \tilde{\mu}_L \rightarrow \tilde{\mu}_R$$

Chiral symmetry would forbid  $g - 2$

# $g - 2$ in the MSSM

chiral symmetry broken by  $\lambda_\mu$ ,  $m_\mu = \lambda_\mu \langle H_1 \rangle$

- second Higgs doublet  $H_2$  important

$$\tan \beta = \frac{\langle H_2 \rangle}{\langle H_1 \rangle}, \quad \mu = H_2 - H_1 \text{ transition}$$

some terms

$$\propto \lambda_\mu \langle H_1 \rangle = m_\mu \quad \rightarrow \quad a_\mu^{\text{SUSY}} \propto \frac{m_\mu^2}{M_{\text{SUSY}}^2}$$

some terms

$$\propto \lambda_\mu \mu \langle H_2 \rangle = m_\mu \mu \tan \beta \quad \rightarrow \quad a_\mu^{\text{SUSY}} \propto \tan \beta \text{ sign}(\mu) \frac{m_\mu^2}{M_{\text{SUSY}}^2}$$

potential enhancement  $\propto \tan \beta = 1 \dots 50$

# $a_\mu$ in the MSSM

1-Loop result if  $\mu, m_{\tilde{\mu}}, m_{\tilde{\chi}} \approx M_{\text{SUSY}}$

$$a_\mu^{\text{SUSY}} \approx \frac{\alpha}{\pi 8s_W^2} \tan \beta \text{sign}(\mu) \frac{m_\mu^2}{M_{\text{SUSY}}^2}$$

numerically

$$a_\mu^{\text{SUSY}} \approx 12 \times 10^{-10} \tan \beta \text{sign}(\mu) \left( \frac{100\text{GeV}}{M_{\text{SUSY}}} \right)^2$$

- $\propto \tan \beta \text{sign}(\mu)$
- $\propto 1/M_{\text{SUSY}}^2$ , but complicated dependence on individual masses

# $a_\mu$ in the MSSM

1-Loop result if  $\mu, m_{\tilde{\mu}}, m_{\tilde{\chi}} \approx M_{\text{SUSY}}$

$$a_\mu^{\text{SUSY}} \approx \frac{\alpha}{\pi 8s_W^2} \tan \beta \text{sign}(\mu) \frac{m_\mu^2}{M_{\text{SUSY}}^2}$$

numerically

$$a_\mu^{\text{SUSY}} \approx 12 \times 10^{-10} \tan \beta \text{sign}(\mu) \left( \frac{100 \text{ GeV}}{M_{\text{SUSY}}} \right)^2$$

e.g.  $a_\mu^{\text{SUSY}} = 24 \times 10^{-10}$  for

$$\begin{aligned} \tan \beta = 2, & \quad M_{\text{SUSY}} = 100 \text{ GeV} \\ \tan \beta = 50, & \quad M_{\text{SUSY}} = 500 \text{ GeV} \end{aligned} \quad (\mu > 0)$$

⇒ SUSY could easily be the origin of the observed deviation!

# $a_\mu$ in the MSSM

1-Loop result if  $\mu, m_{\tilde{\mu}}, m_{\tilde{\chi}} \approx M_{\text{SUSY}}$

$$a_\mu^{\text{SUSY}} \approx \frac{\alpha}{\pi 8s_W^2} \tan\beta \text{sign}(\mu) \frac{m_\mu^2}{M_{\text{SUSY}}^2}$$

numerically

$$a_\mu^{\text{SUSY}} \approx 12 \times 10^{-10} \tan\beta \text{sign}(\mu) \left( \frac{100\text{GeV}}{M_{\text{SUSY}}} \right)^2$$

e.g.  $a_\mu^{\text{SUSY}} = -96 \times 10^{-10}$  for

$$\tan\beta = 50, \quad M_{\text{SUSY}} = 250 \text{ GeV} \quad (\mu < 0)$$

⇒ such parameter points are ruled out by  $a_\mu$ !

# $a_\mu$ in the MSSM

Answers:

SUSY could be the origin of the observed  $(28 \pm 8) \times 10^{-10}$  deviation!

$a_\mu$  significantly restricts the SUSY parameters

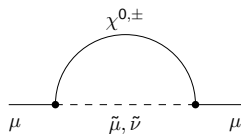
→ generically, positive  $\mu$ , large  $\tan \beta$ /small  $M_{\text{SUSY}}$  preferred

Precise analysis justified!

# Status of SUSY prediction

1-Loop

$$\propto \tan \beta$$



[Fayet '80],...

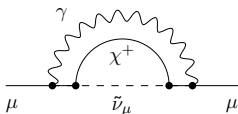
[Kosower et al '83],[Yuan et al '84],...

[Lopez et al '94],[Moroi '96]

complete

2-Loop (SUSY 1L)

$$\propto \log \frac{M_{\text{SUSY}}}{m_\mu}$$

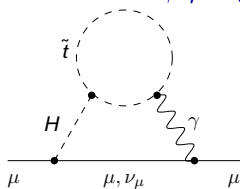


[Degrassi, Giudice '98]

leading log

2-Loop (SM 1L)

$$\propto \tan \beta \mu m_t$$



[Chen, Geng'01][Arhrib, Baek '02]

[Heinemeyer, DS, Weiglein '03]

[Heinemeyer, DS, Weiglein '04]

complete



# SUSY prediction

Implementation of 1-Loop and leading 2-Loop straightforward:

1-Loop: 
$$a_\mu^{\chi^0} = \frac{m_\mu}{16\pi^2} \sum_{i,m} \left\{ -\frac{m_\mu}{12m_{\tilde{\mu}m}^2} (|n_{im}^L|^2 + |n_{im}^R|^2) F_1^N(x_{im}) + \frac{m_{\chi_i^0}}{3m_{\tilde{\mu}m}^2} \text{Re}[n_{im}^L n_{im}^R] F_2^N(x_{im}) \right\},$$

$$a_\mu^{\chi^\pm} = \frac{m_\mu}{16\pi^2} \sum_k \left\{ \frac{m_\mu}{12m_{\tilde{\nu}\mu}^2} (|c_k^L|^2 + |c_k^R|^2) F_1^C(x_k) + \frac{2m_{\chi_k^\pm}}{3m_{\tilde{\nu}\mu}^2} \text{Re}[c_k^L c_k^R] F_2^C(x_k) \right\},$$

$$n_{im}^L = \frac{1}{\sqrt{2}} (g_1 N_{i1} + g_2 N_{i2}) U_{m1}^{\tilde{\mu}*} - y_\mu N_{i3} U_{m2}^{\tilde{\mu}*},$$

$$n_{im}^R = \sqrt{2} g_1 N_{i1} U_{m2}^{\tilde{\mu}} + y_\mu N_{i3} U_{m1}^{\tilde{\mu}},$$

$$c_k^L = -g_2 V_{k1},$$

$$c_k^R = y_\mu U_{k2}. F_1^N(x) = \frac{2}{(1-x)^2} [1 - 6x + 3x^2 + 2x^3 - 6x^2 \log x],$$

$$F_2^N(x) = \frac{3}{(1-x)^3} [1 - x^2 + 2x \log x],$$

$$F_1^C(x) = \frac{2}{(1-x)^2} [2 + 3x - 6x^2 + x^3 + 6x \log x],$$

$$F_2^C(x) = \frac{3}{(1-x)^3} [-3 + 4x - x^2 - 2 \log x],$$

# SUSY prediction

Implementation of 1-Loop and leading 2-Loop straightforward:

2-Loop: 
$$a_\mu^{\text{logs}} = -\frac{4\alpha}{\pi} \log \frac{M_{\text{SUSY}}}{m_\mu} a_\mu^{\text{1-Loop}}$$

$$a_\mu^{(\chi\gamma H)} = \frac{\alpha^2 m_\mu^2}{8\pi^2 M_W^2 S_W^2} \sum_{k=1,2} \left[ \text{Re}[\lambda_\mu^{A^0} \lambda_{\chi_k^+}^{A^0}] f_{PS}(m_{\chi_k^+}^2/M_{A^0}^2) + \sum_{S=H^0, H^0} \text{Re}[\lambda_\mu^{S\lambda_{\chi_k^+}^S}] f_S(m_{\chi_k^+}^2/M_S^2) \right],$$

$$a_\mu^{(\bar{l}\gamma H)} = \frac{\alpha^2 m_\mu^2}{8\pi^2 M_W^2 S_W^2} \sum_{\bar{l}=\bar{t}, \bar{b}, \bar{\tau}} \sum_{i=1,2} \left[ \sum_{S=H^0, H^0} (N_C Q^2)_{\bar{l}} \text{Re}[\lambda_\mu^{S\lambda_{\bar{l}}^S}] f_{\bar{l}}(m_{\bar{l}}^2/M_S^2) \right].$$

$$\lambda_\mu^{\{H^0, H^0, A^0\}} = \left\{ -\frac{s_\alpha}{c_\beta}, \frac{c_\alpha}{c_\beta}, t_\beta \right\},$$

$$\lambda_{\chi_k^+}^{\{H^0, H^0, A^0\}} = \frac{\sqrt{2}M_W}{m_{\chi_k^+}} (U_{k1} V_{k2} \{c_\alpha, s_\alpha, -c_\beta\} + U_{k2} V_{k1} \{-s_\alpha, c_\alpha, -s_\beta\}).$$

$$\lambda_{\bar{l}}^{\{H^0, H^0\}} = \frac{2m_l}{m_{\bar{l}}^2 c_\beta} (+\mu^* \{s_\alpha, -c_\alpha\} + A_l \{c_\alpha, s_\alpha\}) (U_{\bar{l}1}^{\bar{l}})^* U_{\bar{l}2}^{\bar{l}},$$

$$\lambda_{\bar{b}}^{\{H^0, H^0\}} = \frac{2m_b}{m_{\bar{b}}^2 c_\beta} (-\mu^* \{c_\alpha, s_\alpha\} + A_b \{-s_\alpha, c_\alpha\}) (U_{\bar{b}1}^{\bar{b}})^* U_{\bar{b}2}^{\bar{b}},$$

$$\lambda_{\bar{\tau}}^{\{H^0, H^0\}} = \frac{2m_\tau}{m_{\bar{\tau}}^2 c_\beta} (-\mu^* \{c_\alpha, s_\alpha\} + A_\tau \{-s_\alpha, c_\alpha\}) (U_{\bar{\tau}1}^{\bar{\tau}})^* U_{\bar{\tau}2}^{\bar{\tau}}.$$

$$f_{PS}(z) = \frac{2z}{y} \left[ \text{Li}_2\left(1 - \frac{1-y}{2z}\right) - \text{Li}_2\left(1 - \frac{1+y}{2z}\right) \right]$$

$$f_S(z) = (2z - 1)f_{PS}(z) - 2z(2 + \log z),$$

$$f_{\bar{l}}(z) = \frac{z}{2} [2 + \log z - f_{PS}(z)].$$

# SUSY prediction

- 1-loop and most 2-loop contributions known
- remaining theory uncertainty of SUSY prediction: [DS '06]

$$\delta a_\mu^{\text{SUSY}} \approx 3 \times 10^{-10}$$

while

$$\delta a_\mu^{\text{exp-SM}} \approx 8 \times 10^{-10}$$

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# Numerical results

Generic behaviour understood.

Now study details → interesting aspects:

- “typical” / “most general” results
- “aggressive” / “conservative” bounds on SUSY parameters

# Example for “typical” behaviour

## benchmark point SPS1a

$$a_{\mu}(\text{SUSY, SPS1a}) = 29.8(3.1) \times 10^{-10} \quad [\text{DS '06}]$$

$$a_{\mu}(\text{Exp.} - \text{SM}) =$$

$$M_W(\text{SPS1a}) =$$

$$M_W(\text{Exp.}) =$$

## Agreement with experiment?

# Example for “typical” behaviour

## benchmark point SPS1a

$$a_{\mu}(\text{SUSY, SPS1a}) = 29.8(3.1) \times 10^{-10} \quad [\text{DS '06}]$$

$$a_{\mu}(\text{Exp.} - \text{SM}) = 28.7(9.1) \times 10^{-10} \quad [\text{Jegerlehner '07}]$$

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$$M_W(\text{SPS1a}) = 80.381(18) \text{ GeV} \quad [\text{Heinemeyer, Hollik, DS, Weber, Weiglein '06}]$$

$$M_W(\text{Exp.}) =$$

## Agreement with experiment?



# Example for “typical” behaviour

## benchmark point SPS1a

$$a_\mu(\text{SUSY, SPS1a}) = 29.8(3.1) \times 10^{-10} \quad [\text{DS '06}]$$

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$$M_W(\text{SPS1a}) = 80.381(18) \text{ GeV} \quad [\text{Heinemeyer, Hollik, DS, Weber, Weiglein '06}]$$

$$M_W(\text{Exp.}) = 80.398(25) \text{ GeV} \quad [\text{CDF, LEPWWG '07}]$$

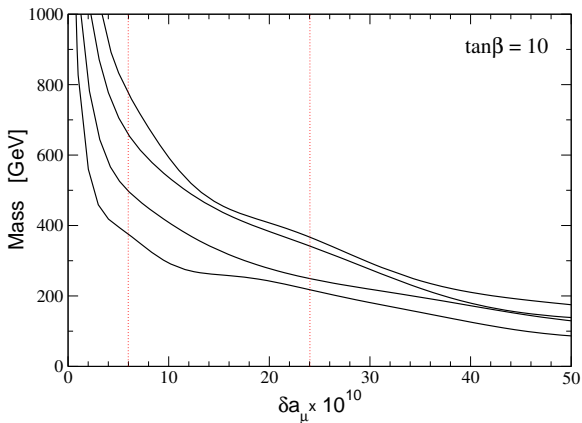
Agreement with experiment? **very good!**

## Numerical results

SPS Point	$a_{\mu}^{\text{SUSY},1\text{L}}(\text{improved})$	$a_{\mu}^{(\chi\gamma H)}$	$a_{\mu}^{(\tilde{f}\gamma H)}$	$a_{\mu}^{\text{SUSY},\chi+\tilde{f},\text{rest}}$	$a_{\mu}^{\text{SUSY},\text{ferm}+\text{bos},2\text{L}}$
SPS 1a	29.29	0.168	0.029	0.056	0.267
SPS 1b	31.84	0.273	0.044	0.106	0.222
SPS 2	01.65	0.032	-0.002	0.027	0.068
SPS 3	13.55	0.078	0.009	0.029	0.187
SPS 4	49.04	0.786	0.085	0.288	0.349
SPS 5	08.59	0.029	0.135	-0.046	0.153
SPS 6	16.87	0.125	0.015	0.044	0.230
SPS 7	23.71	0.236	0	0.089	0.282
SPS 8	17.33	0.163	-0.001	0.062	0.211
SPS 9	-08.98	-0.046	-0.002	-0.018	0.115

[DS '06]

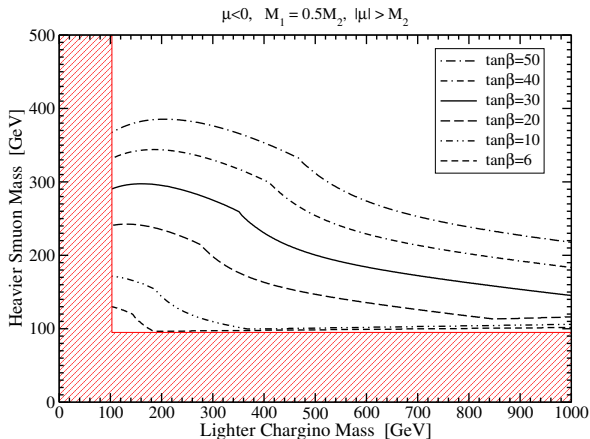
# Numerical results



“aggressive”: require  $a_\mu^{\text{SUSY}}$  within  $2\sigma$  band [Byrne,Kolda,Lennon '02]

⇒ upper mass bounds on four lightest sparticles

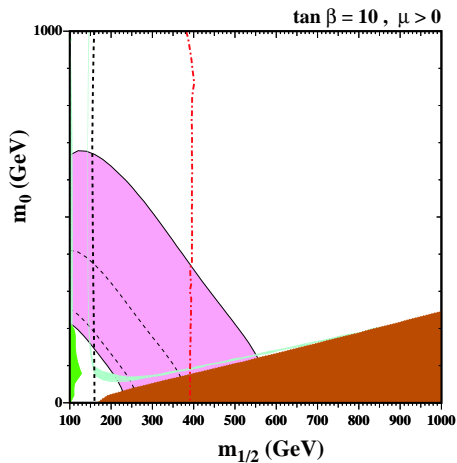
# Numerical results



“conservative”: require  $a_{\mu}^{\text{SUSY}}$  within  $5\sigma$  band [Martin Wells '02]

$\Rightarrow$  lower mass limits

# Numerical results



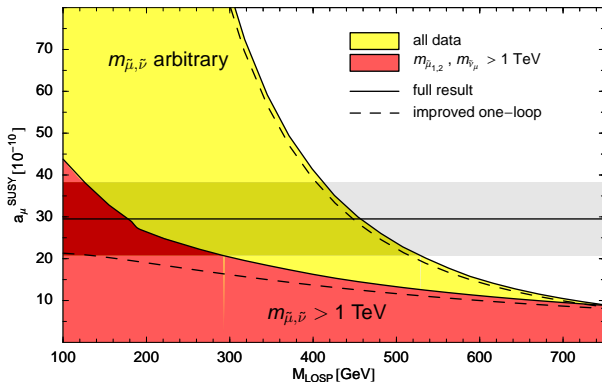
[Ellis, Olive, Sandick '06]

- MSugra scenario ( $\rightarrow$  only two parameters)
- even in very restricted scenarios, SUSY can accommodate the observed value of  $a_\mu$  consistently with many other constraints from dark matter,  $b$ -decays, direct SUSY+Higgs searches

# Numerical results

Summary: scan for  $\tan\beta = 50$ , all parameters  $< 3$  TeV

[DS '06]



- typically:
 
$$12 \times 10^{-10} \tan\beta \times \left(\frac{100 \text{ GeV}}{M_{\text{SUSY}}}\right)^2 \text{sign}(\mu)$$

SUSY contributions in the observed range for low  $M_{\text{SUSY}}$ !

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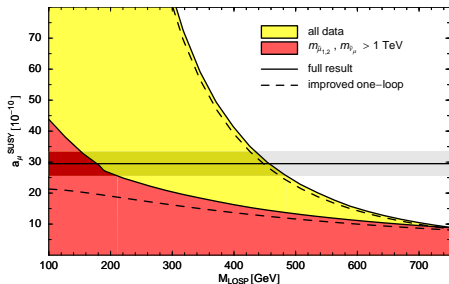
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# Potential of improved measurement

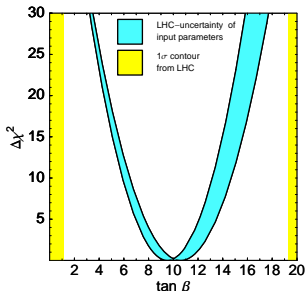
- new Brookhaven experiment proposed and feasible
- improved SM evaluation possible
- projected accuracy:  $a_\mu(\text{Exp-SM}) = 29.5(3.9) \times 10^{-10}$  [Roberts et al 07]

Would be of tremendous importance as a complement of LHC

Constrain SUSY



Measure  $\tan \beta$  (case SPS1a)





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# Conclusions I

- history: fantastic experiments & calculations, many errors. . .
- experiment finalized, SM prediction has recently improved (and will further improve!)

$$a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM, HMNT, DEHZ}} = (28 \pm 8) \times 10^{-10} \quad 3.4\sigma$$

- significance of deviation gets stronger!
- further improvement of hadronic contributions can be expected!

# Conclusions II

$$a_{\mu}^{\text{exp}} - a_{\mu}^{\text{SM, HMNT, DEHZ}} = (28 \pm 8) \times 10^{-10} \quad 3.4\sigma$$

- Case for new physics gets stronger!
- current measurements sensitive to 2-Loop SM and SUSY effects
- 2-Loop SUSY contributions:
  - reliable theory prediction
- SUSY with low mass scale  $\sim 200 \dots 600$  GeV fits very well and large parameter regions already excluded

[Degrassi, Giudice '98]

[Heinemeyer, DS, Weiglein '03,'04]

[DS '06]

Future, more precise measurements very important and promising!